

# The Light Field

---

**Light field = radiance function on rays**

**Conservation of radiance**

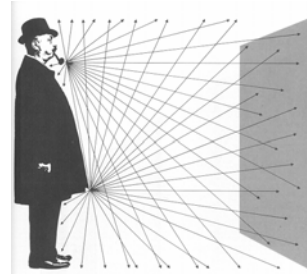
**Measurement equation**

**Throughput and counting rays**

**Conservation of throughput**

**Area sources and irradiance**

**Form factors and radiosity**



**From London and Upton**

CS348B Lecture 5

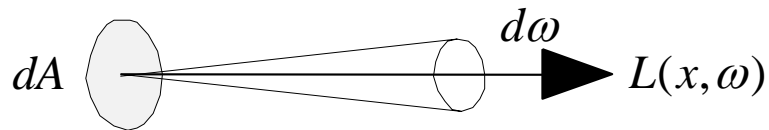
Pat Hanrahan, 2005

**Light Field = Radiance(Ray)**

## Field Radiance

---

**Definition:** The field *radiance* (*luminance*) at a point in space in a given direction is the power per unit solid angle per unit area perpendicular to the direction



**Radiance is the quantity associated with a ray**

CS348B Lecture 5

Pat Hanrahan, 2005

## Light Probe $\Rightarrow$ Environment Map

---

$$L(x, y, z, \theta, \phi)$$



**Miller and Hoffman, 1984**

CS348B Lecture 5

Pat Hanrahan, 2005

# Properties of Radiance

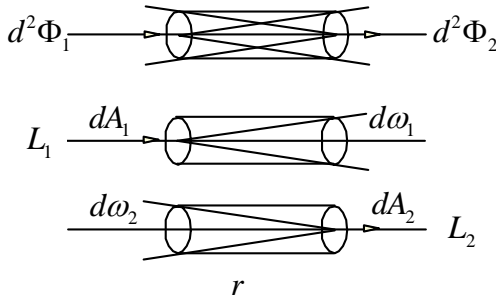
## Properties of Radiance

---

- 1. Fundamental field quantity that characterizes the distribution of light in an environment.**
  - ∴ Radiance is a function on rays
  - ∴ All other field quantities are derived from it
- 2. Radiance invariant along a ray.**
  - ∴ 5D ray space reduces to 4D
- 3. Response of a sensor proportional to radiance.**

# 1st Law: Conversation of Radiance

The radiance in the direction of a light ray remains constant as the ray propagates



$$d^2\Phi_1 = d^2\Phi_2$$

$$d^2\Phi_1 = L_1 d\omega_1 dA_1$$

$$d^2\Phi_2 = L_2 d\omega_2 dA_2$$

$$d\omega_1 dA_1 = \frac{dA_1 dA_2}{r^2} = d\omega_2 dA_2$$

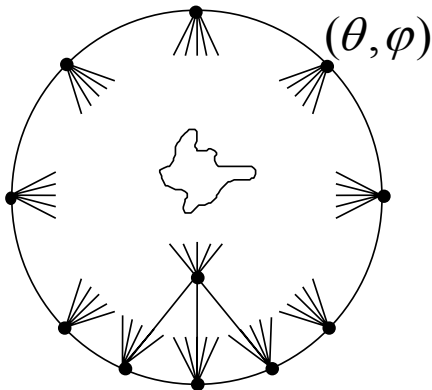
$$\therefore L_1 = L_2$$

CS348B Lecture 5

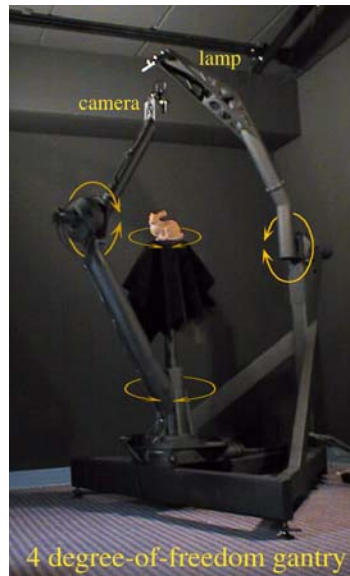
Pat Hanrahan, 2005

# Spherical Gantry $\Rightarrow$ 4D Light Field

$$L(x, y, \theta, \varphi)$$



Capture all the light leaving an object - like a hologram

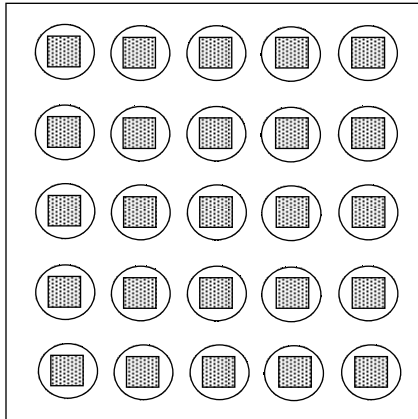


4 degree-of-freedom gantry

CS348B Lecture 5

Pat Hanrahan, 2005

## Two-Plane Light Field



2D Array of Cameras



2D Array of Images

$$L(u, v, s, t)$$

CS348B Lecture 5

Pat Hanrahan, 2005

## Multi-Camera Array $\Rightarrow$ Light Field



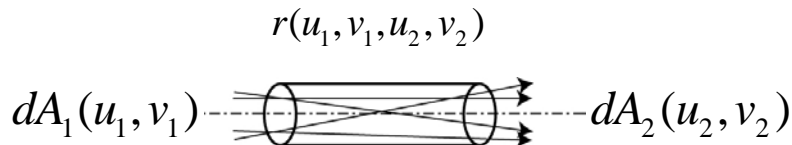
CS348B Lecture 5

Pat Hanrahan, 2005

## Throughput Counts Rays

---

Define an infinitesimal beam as the set of rays intersecting two infinitesimal surface elements



$$d^2T = \frac{dA_1 dA_2}{|x_1 - x_2|^2}$$

**$T$  measures/count the number of rays in the beam**

CS348B Lecture 5

Pat Hanrahan, 2005

## Conservation of Throughput

---

- **Throughput conserved during propagation**
  - **Number of rays conserved**
  - **Assuming no attenuation or scattering**
- **$n^2$  (index of refraction) times throughput invariant under the laws of geometric optics**
  - **Reflection at an interface**
  - **Refraction at an interface**
    - **Causes rays to bend (kink)**
  - **Continuously varying index of refraction**
    - **Causes rays to curve; mirages**

CS348B Lecture 5

Pat Hanrahan, 2005

## Conservation of Radiance

---

Radiance is the ratio of two quantities:

1. Power
2. Throughput

$$L(r) = \lim_{\Delta T \rightarrow 0} \frac{\Delta\Phi(\Delta T)}{\Delta T} = \frac{d\Phi}{dT}$$

Since power and throughput are conserved,

$\therefore$  Radiance conserved

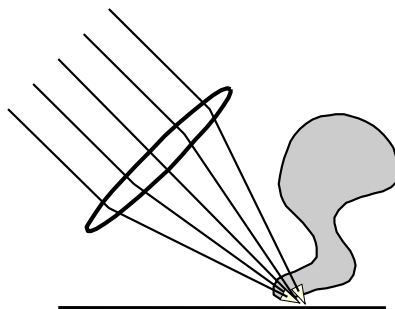
CS348B Lecture 5

Pat Hanrahan, 2005

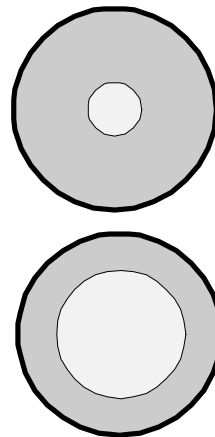
## Quiz

---

Does radiance increase under a magnifying glass?



**No!!**



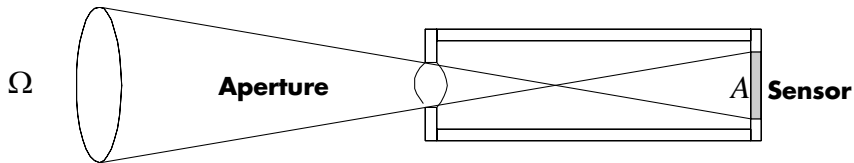
CS348B Lecture 5

Pat Hanrahan, 2005

## Radiance: 2nd Law

---

The response of a sensor is proportional to the radiance of the surface visible to the sensor.



$$R = \int \int_{A \Omega} L d\omega dA = \bar{L}T \quad T = \int \int_{A \Omega} d\omega dA$$

**L** is what should be computed and displayed.

**T** quantifies the gathering power of the device; the higher the throughput the greater the amount of light gathered

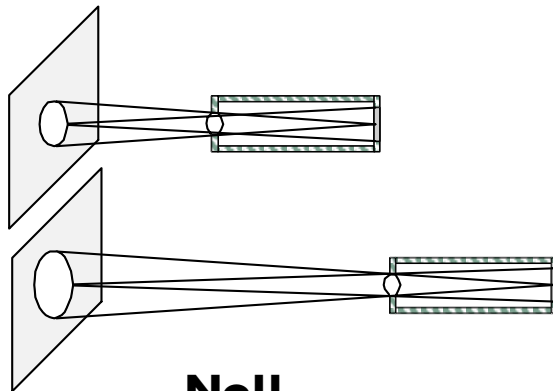
CS348B Lecture 5

Pat Hanrahan, 2005

## Quiz

---

Does the brightness that a wall appears to the sensor depend on the distance?



**No!!**

CS348B Lecture 5

Pat Hanrahan, 2005

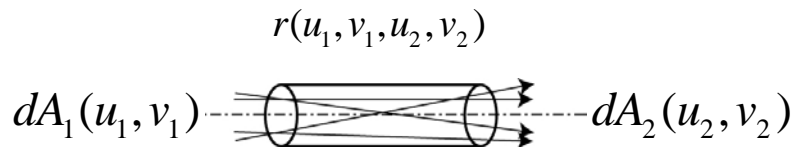


## Measuring Rays = Throughput

### Throughput Counts Rays

---

Define an infinitesimal beam as the set of rays intersecting two infinitesimal surface elements



$$d^2T = \frac{dA_1 dA_2}{|x_1 - x_2|^2}$$

Measure/count the number of rays in the beam

## Parameterizing Rays

---

Parameterize rays wrt to receiver  $r(u_2, v_2, \theta_2, \phi_2)$



$$d^2T = \frac{dA_1}{|x_1 - x_2|^2} dA_2 = d\omega_2 dA_2$$

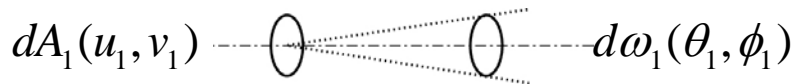
CS348B Lecture 5

Pat Hanrahan, 2005

## Parameterizing Rays

---

Parameterize rays wrt to source  $r(u_1, v_1, \theta_1, \phi_1)$



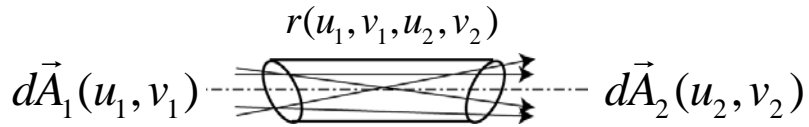
$$d^2T = dA_1 \frac{dA_2}{|x_1 - x_2|^2} = dA_1 d\omega_1$$

CS348B Lecture 5

Pat Hanrahan, 2005

## Parameterizing Rays

Tilting the surfaces reparameterizes the rays



$$d^2T = \frac{\cos \theta_1 \cos \theta_2}{|x_1 - x_2|^2} dA_1 dA_2$$

$$= d\vec{\omega}_1 \cdot d\vec{A}_1$$

$$= d\vec{\omega}_2 \cdot d\vec{A}_2$$

All these throughputs must be equal.

CS348B Lecture 5

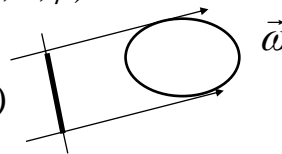
Pat Hanrahan, 2005

## Parameterizing Rays: $S^2 \times R^2$

Parameterize rays by  $r(x, y, \theta, \phi)$

Projected area

$$\tilde{A}(\vec{\omega})$$



Measuring the number or rays that hit a shape

$$T = \int_{S^2} d\omega(\theta, \phi) \int_{R^2} dA(x, y)$$

$$= \int_{S^2} \tilde{A}(\theta, \phi) d\omega(\theta, \phi)$$

$$= 4\pi \tilde{A}$$

**Sphere:**

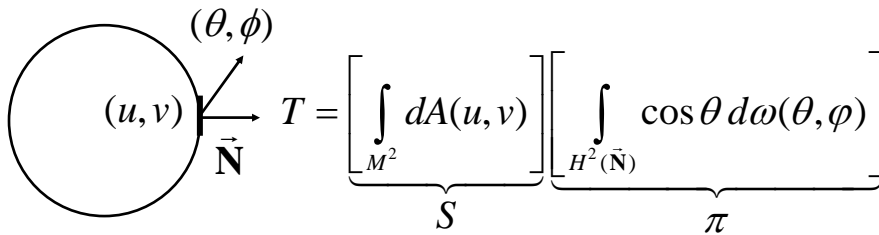
$$T = 4\pi \tilde{A} = 4\pi^2 R^2$$

CS348B Lecture 5

Pat Hanrahan, 2005

## Parameterizing Rays: $M^2 \times S^2$

Parameterize rays by  $r(u, v, \theta, \phi)$



$$T = \underbrace{\left[ \int_{M^2} dA(u, v) \right]}_S \underbrace{\left[ \int_{H^2(\vec{N})} \cos \theta d\omega(\theta, \phi) \right]}_\pi$$

**Sphere:**  $T = \pi S = 4\pi^2 R^2$

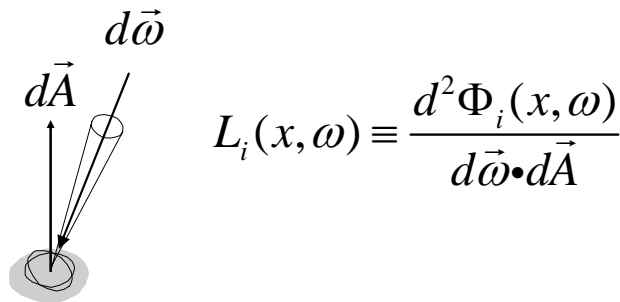
**Crofton's Theorem:**  $4\pi \tilde{A} = \pi S \Rightarrow \tilde{A} = \frac{S}{4}$

CS348B Lecture 5

Pat Hanrahan, 2005

## Incident Surface Radiance

**Definition:** The incoming surface *radiance* (*luminance*) is the power per unit solid angle per unit projected area arriving at a receiving surface



$$L_i(x, \omega) \equiv \frac{d^2 \Phi_i(x, \omega)}{d\vec{\omega} \cdot d\vec{A}}$$

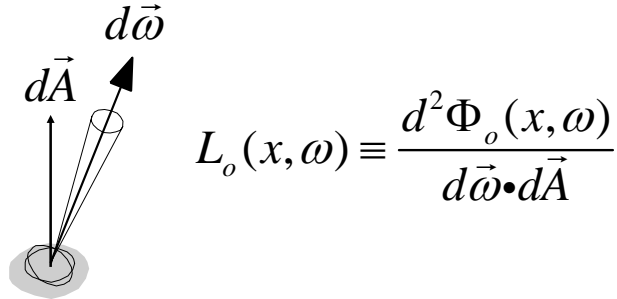
CS348B Lecture 5

Pat Hanrahan, 2005

## Exitant Surface Radiance

---

**Definition:** The outgoing surface radiance (luminance) is the power per unit solid angle per unit projected area leaving at surface



**Alternatively:** the intensity per unit projected area leaving a surface

CS348B Lecture 5

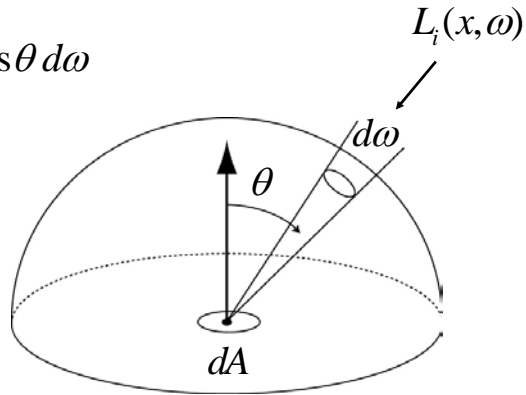
Pat Hanrahan, 2005

## Irradiance from a Uniform Area Source

## Irradiance from the Environment

$$d^2\Phi_i(x, \omega) = L_i(x, \omega) \cos \theta dA d\omega$$

$$dE(x, \omega) = L_i(x, \omega) \cos \theta d\omega$$



$$E(x) = \int_{H^2} L_i(x, \omega) \cos \theta d\omega$$

CS348B Lecture 5

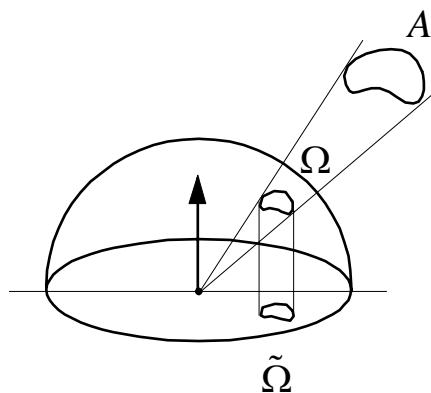
Pat Hanrahan, 2005

## Uniform Area Source

$$E(x) = \int_{H^2} L \cos \theta d\omega$$

$$= L \int_{\Omega} \cos \theta d\omega$$

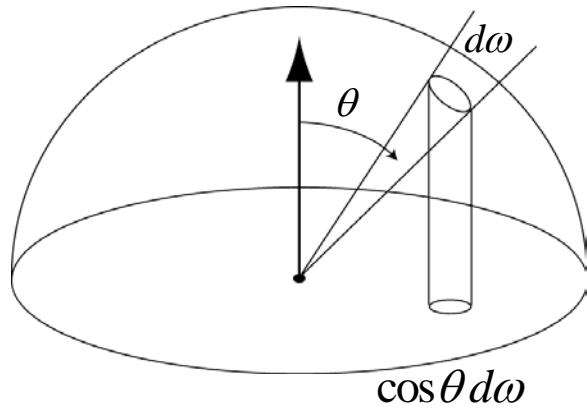
$$= L \tilde{\Omega}$$



CS348B Lecture 5

Pat Hanrahan, 2005

## Projected Solid Angle



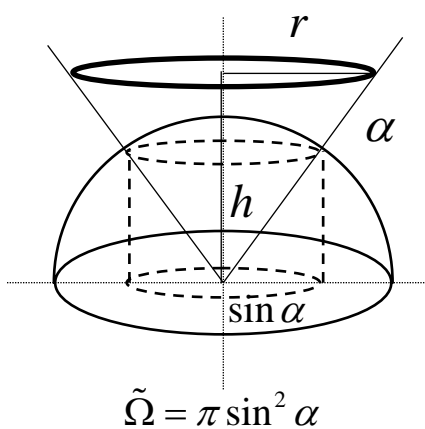
$$\int_{H^2} \cos \theta d\omega = \pi$$

CS348B Lecture 5

Pat Hanrahan, 2005

## Uniform Disk Source

### Geometric Derivation



$$\tilde{\Omega} = \pi \sin^2 \alpha$$

### Algebraic Derivation

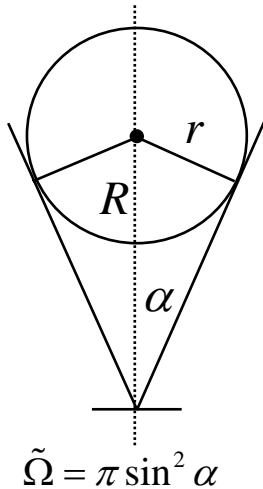
$$\begin{aligned} \tilde{\Omega} &= \int_1^{\cos \alpha} \int_0^{2\pi} \cos \theta d\phi d \cos \theta \\ &= 2\pi \frac{\cos^2 \theta}{2} \Big|_1^{\cos \alpha} \\ &= \pi \sin^2 \alpha \\ &= \pi \frac{r^2}{r^2 + h^2} \end{aligned}$$

CS348B Lecture 5

Pat Hanrahan, 2005

## Spherical Source

### Geometric Derivation



### Algebraic Derivation

$$\begin{aligned}\tilde{\Omega} &= \int \cos \theta d\omega \\ &= \pi \sin^2 \alpha \\ &= \pi \frac{r^2}{R^2}\end{aligned}$$

CS348B Lecture 5

Pat Hanrahan, 2005

## The Sun

### Solar constant (normal incidence at zenith)

**Irradiance**     **1353 W/m<sup>2</sup>**

**Illuminance**     **127,500 lm/m<sup>2</sup> = 127.5 kilolux**

### Solar angle

$\alpha = .25$  degrees = **.004 radians (half angle)**

$\tilde{\Omega} = \pi \sin^2 \alpha \approx \pi \alpha^2 = \mathbf{6 \times 10^{-5}}$  steradians

### Solar radiance

$$L = \frac{E}{\tilde{\Omega}} = \frac{1.353 \times 10^3 \text{ W/m}^2}{6 \times 10^{-5} \text{ sr}} = 2.25 \times 10^7 \frac{\text{W}}{\text{m}^2 \cdot \text{sr}}$$

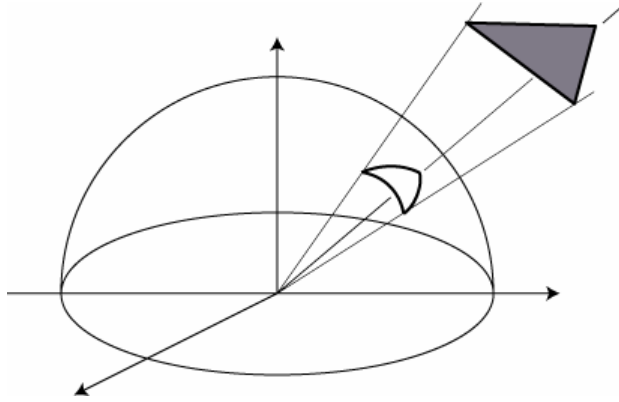
CS348B Lecture 5

Pat Hanrahan, 2005



# Polygonal Source

---

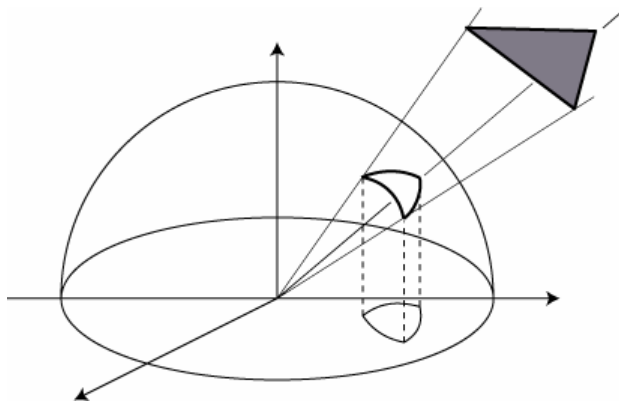


CS348B Lecture 5

Pat Hanrahan, 2005

# Polygonal Source

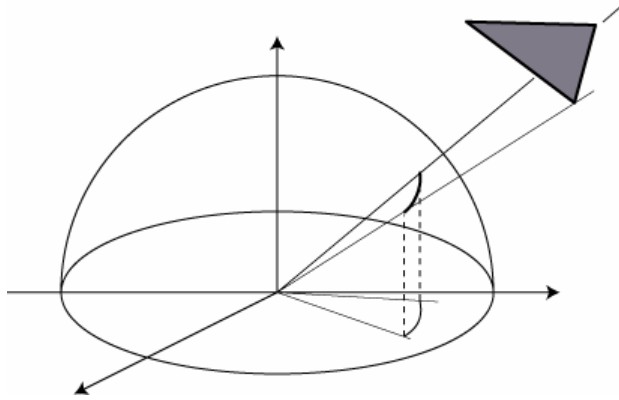
---



CS348B Lecture 5

Pat Hanrahan, 2005

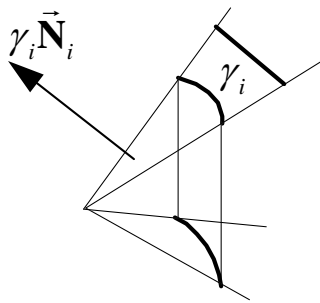
# Polygonal Source



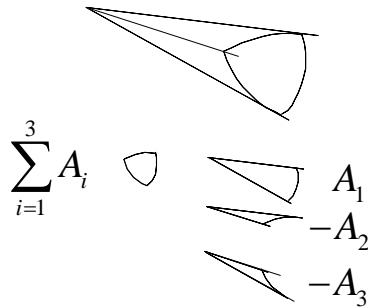
CS348B Lecture 5

Pat Hanrahan, 2005

# Lambert's Formula



$$A_i = \gamma_i \vec{N}_i \cdot \vec{N}$$



$$\sum_{i=1}^3 A_i$$

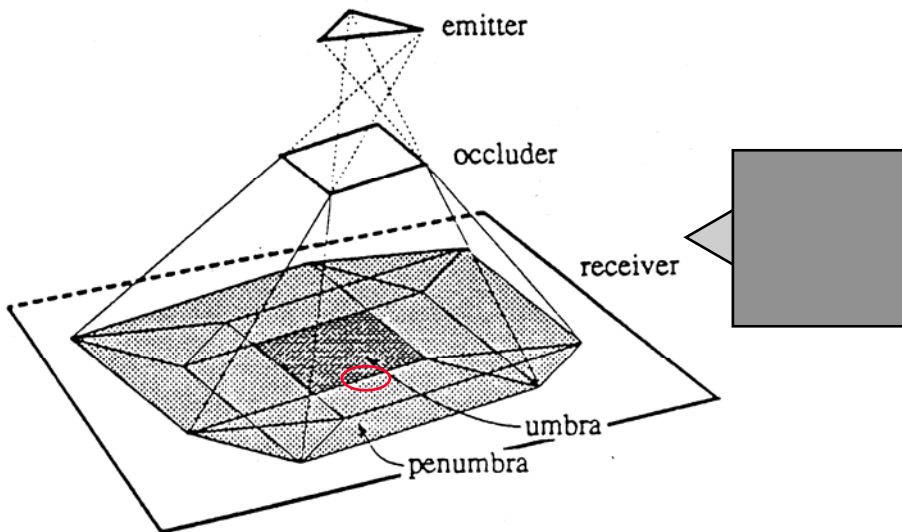
$$\sum_{i=1}^n A_i = \sum_{i=1}^n \gamma_i \vec{N}_i \cdot \vec{N}$$

CS348B Lecture 5

Pat Hanrahan, 2005

## Penumbras and Umbra

---



CS348B Lecture 5

Pat Hanrahan, 2005

## Form Factors

# Types of Throughput

## 1. Infinitesimal beam of rays (radiance)

$$d^2T(dA, dA') \equiv \frac{\cos \theta \cos \theta'}{|x - x'|^2} dA(x) dA(x')$$

## 2. Infinitesimal-finite beam (irradiance calc.)

$$dT(dA, A') dA \equiv \left[ \int_{M'} \frac{\cos \theta \cos \theta'}{|x - x'|^2} dA(x') \right] dA(x)$$

## 3. Finite-finite beam (radiosity calc.)

$$T(A, A') \equiv \iint_{MM'} \frac{\cos \theta \cos \theta'}{|x - x'|^2} dA(x') dA(x)$$

CS348B Lecture 5

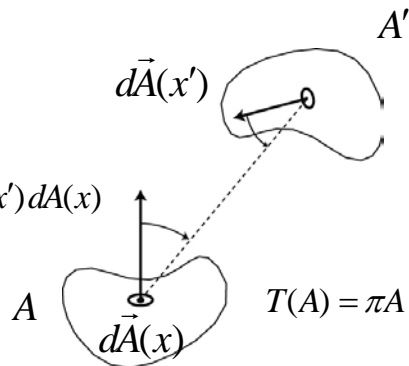
Pat Hanrahan, 2005

# Probability of Ray Intersection

## Probability of a ray hitting $A'$ given that it hits $A$

$$\begin{aligned} \Pr(A' | A) &= \frac{T(A', A)}{T(A)} \\ &= \frac{T(A', A)}{\pi A} \end{aligned}$$

$$T(A', A) = \iint_{MM'} \frac{\cos \theta \cos \theta'}{|x - x'|^2} dA(x') dA(x)$$



CS348B Lecture 5

Pat Hanrahan, 2005

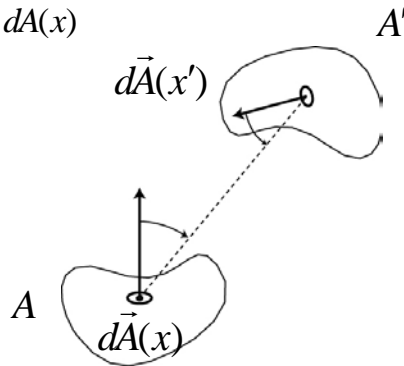
## Another Formulation

$$T(A', A) = \int_{MM'} \int_{MM'} \frac{\cos \theta \cos \theta'}{|x - x'|^2} dA(x') dA(x)$$

$$= \pi \int_{MM'} \int_{MM'} G(x, x') dA(x') dA(x)$$

$$G(x, x') \equiv \frac{\cos \theta \cos \theta'}{\pi |x - x'|^2} V(x, x')$$

$$V(x, x') = \begin{cases} 0 & \text{-visible} \\ 1 & \text{visible} \end{cases}$$



CS348B Lecture 5

Pat Hanrahan, 2005

## Form Factor

### Probability of a ray hitting $A'$ given it hits $A$

$$\Pr(A' | A) T(A) = \Pr(A | A') T(A) = T(A', A)$$

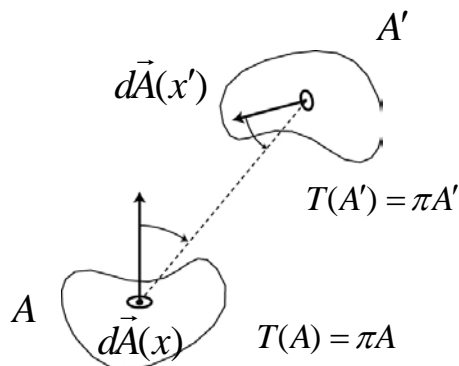
### Form factor definition

$$F(A', A) = \Pr(A' | A)$$

$$F(A, A') = \Pr(A | A')$$

### Form factor reciprocity

$$F(A', A)A = F(A, A')A'$$



CS348B Lecture 5

Pat Hanrahan, 2005

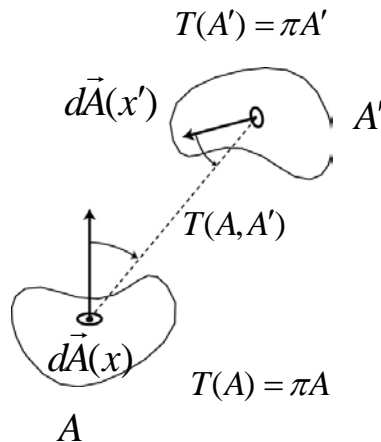
# Radiosity

## Power transfer from a constant radiance source

$$\begin{aligned}\Phi(A, A') &= L(A')T(A, A') \\ &= L(A')T(A') \frac{T(A, A')}{T(A')} \\ &= \Phi(A')F(A, A')\end{aligned}$$

$$E(A) = B(A')F(A', A)$$

**Set up system of equations representing power transfers between objects**



CS348B Lecture 5

Pat Hanrahan, 2005

# Form Factors and Throughput

**Form factors represent the probability of ray leaving a surface intersecting another surface**

- Only a function of surface geometry

**Differential form factor**

- Irradiance calculations

**Form factors**

- Radiosity calculations (energy balance)

CS348B Lecture 5

Pat Hanrahan, 2005